

## Stability of the Nuclear Atom

Rutherford's model  $\Rightarrow$  at the center of the atom there is a nucleus  
 where  $m_{\text{nucleus}} \sim m_{\text{atom}}$  and  $q_{\text{nucleus}} = Ze$   
 around nucleus:  $Z e^-$

but stability?

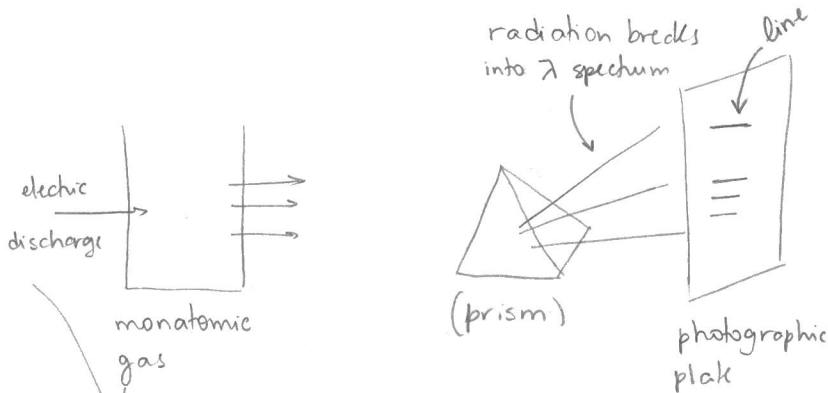
o) stationary  $e^- \Rightarrow$  no possible stable arrangement  
 $e^-$  would fall into nucleus

o) circulating as in the solar system  $\Rightarrow e^-$   
 would lose energy  
 (accelerated  $e^- \Rightarrow$  radiation) and  
 fall into nucleus

ALSO continuous spectrum that would be emitted in the  
 process is not observed, spectrum is discrete

Solution: Bohr's model

## Atomic Spectra



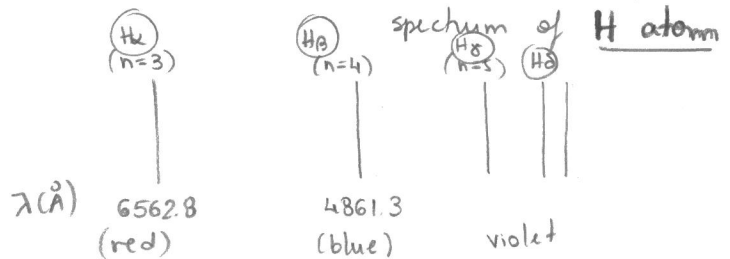
collisions  $\Rightarrow$  some atoms have more energy. When returning to normal energy state  $\rightarrow$  give up excess energy by emitting electromagnetic radiation

electromagnetic radiation emitted by free atoms is concentrated at a number of discrete wavelengths

Each <sup>kind of</sup> atom has its own characteristic spectrum

### SPECTROSCOPY

(1885) Balmer - access to VISIBLE



spacing decreases as  $\lambda$  decreases

$$\lambda = 3646 \frac{n^2}{n^2 - 4}$$

1890 Rydberg: → extended to other elements

$$\text{Balmer formula written for } 1/\lambda \text{ for H} \rightarrow K = \frac{1}{\lambda} = R_H \left( \frac{1}{2^2} - \frac{1}{n^2} \right) \quad n=3,4,5$$

Balmer formula written for  $1/\lambda$  for H

$$\text{Rydberg constant} = 10967757.6 \pm 1.2 \text{ m}^{-1}$$

See Table 4-1, Eisberg, for Hydrogen series for UV and IR

Spectrum of free atoms → discrete

spectroscopy

stars

### Bohr's postulate

↳ explains stability of lines of the spectrum of atoms

- 1)  $e^-$  moves in circular orbit around nucleus due to Coulomb attraction obeying laws of classical mechanics
- 2) only possible orbits: those for which  $L$  is an integral multiple of  $\hbar$ 

$$\hbar = h/2\pi \quad (L = n\hbar)$$

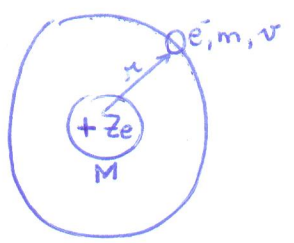
↑ quantum numbers
- 3)  $e^-$  are constantly accelerating, but when moving in allowed orbit they do not radiate,  $E$  is const (stationary states)
- 4) Electromagnetic radiation is emitted only when an  $e^-$  moves from one allowed orbit to another one. The frequency of the emitted radiation is

$$\nu = \frac{E_i - E_f}{h}$$

# Bohr's Model

→ predictions derived from postulates agree with experiments

consider a single  $e^-$



- $Z=1$  (hydrogen) - neutral
- $Z=2$  (singly ionized helium)
- $Z=3$  (doubly ionized lithium)

→ Assume:

- ) circular orbit
- )  $M \gg m \Rightarrow$  nucleus is fixed

$$\textcircled{*} \frac{kZe^2}{r^2} = \frac{mv^2}{r}$$

Coulomb                       $ma \rightarrow$  unbalanced acceleration

→ Orbital angular momentum of  $e^-$ :

$$L = mvr$$

only some radii exist  
only some speeds exist

quantization condition  $L = nh$

$$mvr = nh \quad n = 1, 2, 3, \dots$$

$$\Rightarrow v = \frac{nh}{mr}$$

→ Find RADIUS:  $r = \frac{kZe^2}{mv^2} \Rightarrow$

$$r = \frac{kZe^2}{m \left(\frac{nh}{mr}\right)^2}$$

$$r = \frac{n^2 h^2}{k m Z e^2}$$

$$\begin{aligned} &> Z, < n \\ &> n, > r \end{aligned}$$

$$a_0 = \frac{h^2}{m k e^2} \quad (\text{Bohr radius})$$

$= 0.529 \text{ \AA}$  ← Hydrogen atom in the ground state

$n = 1, 2, 3, \dots$   
 $e^-$  has minimum energy

→ Find speed of  $e^-$ :

$$v = \frac{kZe^2}{nh}$$

$n = 1, 2, 3, \dots$

$\Rightarrow n$  cannot be zero

$$\begin{aligned} > Z, \Rightarrow > v \\ > n, \Rightarrow < v \end{aligned}$$

from Coulomb  $\frac{kZe^2}{r^2} = m \frac{v^2}{r}$  <sup>anticipated</sup>

→ TOTAL energy of  $e^-$

$$K = \frac{1}{2} m v^2 = \frac{kZe^2}{2r}$$

$$U = -\frac{kZe^2}{r} \leftarrow \text{from}$$

$$U = -\int_r^{\infty} \frac{kZe^2}{r^2} dr = -\frac{kZe^2}{r}$$

since Coulomb force is attractive, we need to do work to move  $e^-$  from  $r$  to  $\infty$  against this force

$$E = K + U = -\frac{kZe^2}{2r}$$

$$r = \frac{n^2 \hbar^2}{kmZe^2}$$

need energy to remove  $e^-$

$$E_0 = \frac{k^2 m e^4}{2 \hbar^2}$$

$$E = -\frac{k^2 m (Ze^2)^2}{2 \hbar^2} \frac{1}{n^2}$$

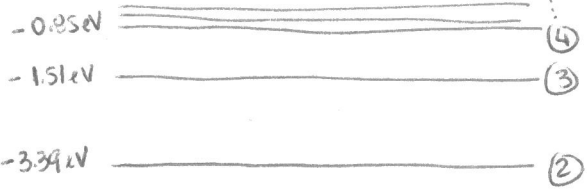
$$n = 1, 2, 3, \dots$$

$$= -\frac{Z^2 E_0}{n^2}$$

→ quantization of  $L$  leads to quantization of total energy of  $e^-$

$E$  (total energy)

$n$  (quantum number)



↑ excited states



lowest energy = ground state

most stable state

most stable state, most negative

need the largest energy to remove  $e^-$

$E_0$  { the binding energy of H is -13.6 eV  
or ionization energy of H is +13.6 eV

→ Frequency of radiation emitted when  $e^-$  makes a transition

from state  $n_i$  to  $n_f$

$$\nu = \frac{E_i - E_f}{h} = \frac{-\frac{k^2 m (Ze^2)^2}{2 \hbar^2} \left( \frac{1}{n_i^2} - \frac{1}{n_f^2} \right)}{h} = \frac{+ \frac{k^2 m (Ze^2)^2}{2 \hbar^2} \left( \frac{1}{n_f^2} - \frac{1}{n_i^2} \right)}{h}$$

$\hookrightarrow h 2\pi$

Electron makes a transition from state  $n_i$  to  $n_f$

$$\nu = \frac{E_i - E_f}{h} = \frac{E_0 Z^2}{h} \left( \frac{1}{n_f^2} - \frac{1}{n_i^2} \right)$$

$$K = 1/\lambda = \nu/c$$

$$K = \frac{1}{\lambda} = \frac{E_0 Z^2}{hc} \left( \frac{1}{n_f^2} - \frac{1}{n_i^2} \right)$$

Bohr's prediction for the Rydberg constant  
 $R_{\infty}$   $R_H$

$n_f = 1$  Lyman series

$n_f = 2$  Balmer series:  $K = R_{\infty} \left( \frac{1}{2^2} - \frac{1}{n_i^2} \right)$

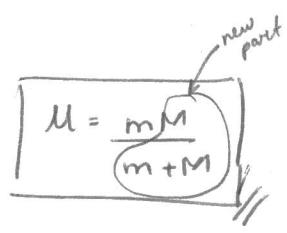
⋮

→ Correction for Finite Nuclear Mass (Bohr's assumption that nucleus is fixed  
 ↳  $M$  is infinite)

↳ substitute  $m$  by  $\mu$ .  $\mu =$  reduced mass

$$\text{TOTAL } K = \frac{p^2}{2M} + \frac{p^2}{2m} = \frac{M+m}{2mM} p^2 = \frac{p^2}{2\mu}$$

nucleus  
 electron  
 nucleus = electron (atom at rest)  
 $p_{nucleus} = 0$



$$R = \frac{k^2 \mu e^4}{4\pi c h^3} = R_{\infty} \left( \frac{1}{1+m/M} \right)$$